

ESS 265

Introduction to XTRANS and
Coordinate Transformation in
Conjunction with Problem Set 4

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Geophysical Coordinate Transformations

- Many different coordinate systems are used to define measurement and observer locations both on Earth and in space.
- Generally, there is a coordinate system that simplifies some aspect of the problem, such as making a variation 2-dimensional.
- It can be very tricky to transform from one system to another, especially when angles and spherical trigonometry are involved.
- Vector-matrix formulation can simplify this process. The XTRANS program is designed to help the user by providing many of these transformations in use in solar terrestrial physics.
- Details of coordinate systems and their usage can be found in Appendix 3 of Kivelson and Russell.

Simple Rules of Coordinate Transformation

- If the *matrix* A transforms *vector* V^a measured in *system* a to V^b measured in *system* b , then its transpose A^T transforms V^b into V^a

$$A \cdot V^a = V^b$$

$$A^T \cdot V^b = V^a$$

- To construct the transformation *matrix* A , find the direction cosines of the three new coordinate axes for *system* b in *system* a coordinates. Let these vectors be (x_1, x_2, x_3) , (y_1, y_2, y_3) and (z_1, z_2, z_3) . Then

$$\begin{pmatrix} x_1 & x_2 & x_3 \\ y_1 & y_2 & y_3 \\ z_1 & z_2 & z_3 \end{pmatrix} \begin{pmatrix} v_x^a \\ v_y^a \\ v_z^a \end{pmatrix} = \begin{pmatrix} v_x^b \\ v_y^b \\ v_z^b \end{pmatrix}$$

and

$$\begin{pmatrix} x_1 & y_1 & z_1 \\ x_2 & y_2 & z_2 \\ x_3 & y_3 & z_3 \end{pmatrix} \begin{pmatrix} v_x^b \\ v_y^b \\ v_z^b \end{pmatrix} = \begin{pmatrix} v_x^a \\ v_y^a \\ v_z^a \end{pmatrix}$$

Commonly Used Coordinate Systems

- Geocentric Equatorial Inertial (GEI)
 - X axis to the first point of Aries (direction of Sun to Earth about September 21)
 - Z axis parallel to Earth's rotation axis
 - $Y = Z \times X$Used for spacecraft orbit calculations
- Geographic Coordinate System (GEO)
 - X axis from center of Earth to Greenwich meridian in equatorial plane
 - Z axis parallel to rotation axis of Earth
 - $Y = Z \times X$Used for terrestrial maps
- GEO differs from GEI by a rotation about Z dependent on sidereal not universal (solar) time (UT).
- Universal Time is defined as 12h minus the longitude of the Sun converted to degrees by dividing by 15.

Geophysical Coordinates

- Geomagnetic coordinates (MAG) have Z along the magnetic dipole axis northward. The Y axis is perpendicular to the plane containing the rotational poles and the dipole. This coordinate system has a fixed transformation from GEO to the extent that the Earth's internal field is steady. This coordinate system is useful for geomagnetic studies at and near the surface of the Earth.
- Geocentric Solar Ecliptic (GSE) has its X axis to the Sun and Z along the north pole of the Earth's orbit.
- Geocentric Solar Equatorial (GSEQ) has its X axis to the Sun with the Sun's rotational axis in the X-Z plane (northward). It orders interplanetary magnetic field data.
- Geocentric Solar Magnetospheric (GSM) has its X axis to the Sun with the Earth's dipole in the X-Z plane. This is used to order observations near the magnetopause and in the magnetotail.
- Solar Magnetic (SM) has its Z axis along the dipole axis and the X-Z plane contains the directions to the Sun. It is used to order measurements in the inner magnetosphere. Magnetic local time is defined in this coordinate system.

Other Coordinate Systems

- Many coordinate systems are defined by the location of the observer.
 - Some coordinates depend on the orientation of a boundary normal
 - Shock normals
 - Magnetopause normals
- These coordinate systems can be derived from models or data.
- The study of waves particularly benefits from the use of the principal axis system.

XTRANS Program

- XTRANS can transform vectors into other coordinate systems, even those that are time-dependent.
- It works in either cartesian or spherical coordinates.
- It provides the transformation matrix.
- It can also determine local time and local magnetic time.



Geographic to Geomagnetic

Geographic[GEO] To Geomagnetic[MAG]

TRANSFORMATION INPUT

Enter Date & Time:

1	1	1994
Month	Day	Year
0	0	0.0
Hour	Minutes	Seconds

Choose & Enter Coordinates:

Spherical:

+1.00000	+0.00000	+0.00000
Magnitude	Longitude(deg)	Latitude(deg)

Cartesian:

+1.00000	+0.00000	+0.00000
X	Y	Z

Press To Compute Output:

TRANSFORMATION OUTPUT

Transformation Matrix:

+0.31453	-0.93071	-0.18668
+0.94736	+0.32016	+0.00000
+0.05977	-0.17685	+0.98242

Transformed Vectors:

Spherical Coords:

+1.00000	+71.63	+3.43
Magnitude	Longitude(deg)	Latitude(deg)

Cartesian Coords:

+0.31453	+0.94736	+0.05977
X	Y	Z

Compute a transformation matrix, or transform a input vector.

GSM to SM

Solar Magnetospheric[GSM] To Solar Magnetic[SM]

TRANSFORMATION INPUT

Enter Date & Time:

5	24	2005
Month	Day	Year
9	30	0.0
Hour	Minutes	Seconds

Choose & Enter Coordinates:

Spherical:

+1.00000	-118.88467	34.188333
Magnitude	Longitude(deg)	Latitude(deg)

Cartesian:

-0.39958	-0.70409	+0.50191
X	Y	Z

Press To Compute Output:

TRANSFORMATION OUTPUT

Transformation Matrix:

+0.95539	+0.00000	-0.29535
+0.00000	+1.00000	+0.00000
+0.29535	+0.00000	+0.95539

Transformed Vectors:

Spherical Coords:

+1.00000	-127.10	+24.76
Magnitude	Longitude(deg)	Latitude(deg)

Cartesian Coords:

-0.54771	-0.72429	+0.41883
X	Y	Z

Compute a transformation matrix, or transform a input vector.

Local Time and Dipole Axis

Compute Local Time & Dipole Axis From Universal Time

OBSERVER'S TIME & POSITION INPUT:
Enter Universal Date & Time:

05	16	2005
Month	Day	Year
19	02	0.0
Hour	Minutes	Seconds

Enter Observers Location (In Geographic):

45.0	45.0
Longitude	Latitude

Press To Select Ground Station, Else Enter Manually Above:

NONE - Selected

Press To Compute Output:

OBSERVER'S TIME & POSITION OUTPUT

Observers Local Time:

22	2	0.00
Hour	Minutes	Seconds

Observers Local Magnetic Time:

22	48	55.24
Hour	Minutes	Seconds

Observers Location (In Geomagnetic):

125.13	39.59
Longitude	Latitude

Location of Dipole Axis:

-71.97	10.43
Longitude	Co-Latitude

Input Universal Time & Observer Location.